

Investigating the datum parameters of new solutions by IVS AC DGFI-TUM

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Solutions investigated

- dgf2020a resembles DGFI-TUM's contribution to the ITRS 2020 realizations (plus reduction of non-tidal atmospheric loading) and contains the respective models (e.g., the gravitational deformation for EFLSBERG, GILCREEK, MEDICINA, NOTO, ONSALA60, YEBES40M).
- dgf2023a is the current operational solution based on a priori values from the new ITRF2020 and containing 7 additional gravitational deformation models.
- To investigate the impact of these two changes, the following (intermediate) setups have been analysed with the DGFI Orbit and Geodetic parameter estimation Software (DOGS):

solution	a priori TRF	gravitational deformation (GD)	datum stations	
dgf2020a	ITRF2014	first 6 antennas	all except for 6 GD antennas	
dgf2023a old GD*	ITRF2020	first 6 antennas	all except for 6 GD antennas	
dgf2023a old GD	ITRF2020	first 6 antennas	all	
dgf2023a*	ITRF2020	first 6 + 7 new antennas	all	
dgf2023a	ITRF2020	first 6 + 7 new antennas	all except for 7 new GD antennas	
dgf2023a JTRF	JTRF2020	first 6 + 7 new antennas	all except for 7 new GD antennas	
dgf2023a DTRF	DTRF2020	first 6 + 7 new antennas	all except for 7 new GD antennas	

ТЛП

ITRS 2020 realizations

➢ ITRF2020

- secular terrestrial reference frame (TRF) with reduction of seasonal (and draconitic) signals and post-seismic deformation (PSD)
- combined at the solution level

JTRF2020

- epoch reference frame
- sequentially determined with a square-root information filter and smoother algorithm

DTRF2020

- secular TRF with reduction of non-tidal loading (NTL) and PSD
- combined at the normal equation level
- NTL based on geophysical models, reduced at normal equation level



Solution **dgf2023a** with ITRF2020 (secular plus PSD) as a priori TRF and 7 new gravitational deformation models. Only the **secular and PSD parts** of both ITRF2020 and DTRF2020 are given for comparison.

Impact of new a priori TRF

solution	a priori TRF	gravitational deformation (GD)	datum stations
dgf2020a	ITRF2014	first 6 antennas	all except for 6 GD antennas
dgf2023a old GD*	ITRF2020	first 6 antennas	all except for 6 GD antennas
dgf2023a old GD	ITRF2020	first 6 antennas	all

- If the other models do not change, a new a priori TRF
 does not change the antenna network geometry.
- However, the no-net-translation (NNT) and no-net-rotation
 (NNR) conditions affect the final coordinate estimates.
- If the set of datum stations remains constant, the impact is comparatively small if the antenna coordinates in a session network do not change much (red dots in figures).
- Changing the set of datum stations creates additional noise (blue dots).
- After 2015, there are significant changes for both choices of datum station sets (compare next slide).



Differences between latest ITRFs

> After 2015, there are significant **changes in linear motions of "young" stations:**



ITRF2020 minus ITRF2014

Each color represents the difference time series for one VLBI station (secular + PSD parts only).

Differences between latest ITRFs (up)



> After 2015, there are significant changes in linear motions of "young" stations:

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Datum (scale) parameter time series: ITRS 2020 realizations

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dgf2023a JTRF	JTRF2020	first 6 + 7 new antennas	all except for 7 new GD antennas	
dgf2023a DTRF	DTRF2020	first 6 + 7 new antennas	all except for 7 new GD antennas	

- The figures show the scale parameter of 7 parameter Helmert transformations of the session-wise solutions w.r.t. their a priori TRFs (outliers: |scale| > 60 mm).
- The running 20 session median reveals an annual signal for the secular TRFs and a similar signal for the JTRF2020.
- There is a scale drift for ITRF2020 and JTRF2020, but none for DTRF2020.
- > Translations and rotations vary around zero.



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- > Translations and rotations vary around zero.
- For ITRF2020 and JTRF2020, basically all scales for the VGOS sessions are positive!



Differences between secular ITRS 2020 realizations (up)





Further models for gravitational deformation (1)

- As of January 2023, there are 7 more antennas with gravitational deformation (GD) models available (compare file compiled by J.Gipson):
 - NYALES20
 - KOKEE
 - ONSA13NE (VGOS)
 - ONSA13SW (VGOS)
 - WETTZELL
 - WETTZ13N
 - WETTZ13S (VGOS)
- > GD mainly affectes the estimated antenna heights, which are proportional to $-\sin \epsilon$.
- Leaving antennas with GD models among the set of (NNT) datum stations transfers vertical to horizontal changes (compare Gipson [2019]).
- When switching to ITRF2020 as a priori TRF, the first
 6 antennas can be considered as datum stations
 again, because their GD models are part of this TRF.



 * YEBES40M: only valid from 2011-11-11
 x KOKEE: values adopted from NYALES20
 # ONSALA60: original values -4.7 mm to adjust reference temperature w.r.t. thermal deformation

New models represented by dashed lines.



Further models for gravitational deformation (2)

- The actual change in estimated antenna heights (see table*) depends on the maximum model excess delay and the sign of its slope (w.r.t. elevation).
- The maximum delays for the 7 new antennas are mostly smaller than those for the 6 old ones.
- We expect an increase in heights for the new Onsala twin telescopes only, and a decrease for the other 5 antennas.
- However, dropping antennas with long observation histories and/or remote locations (like KOKEE, NYALES20, WETTZELL) from the datum set is little desirable.
- > What is the **best approach**?
- * J.Gipson, Impact of Gravitational Deformation of VLBI Antennas on Reference Frame, UAW 2019;
- M.Glomsda et al., Impact of new models for the ITRF2020 in VLBI analysis at DGFI-TUM, AGU 2020



	EFLSBERG	GILCREEK	MEDICINA	ΝΟΤΟ	ONSALA60	YEBES40M
GSFC estimated Δh [mm]	118.70	2.46	8.87	7.26	-4.92	-37.44
DGFI-TUM wmean Δh [mm]	119.1	2.4	9.2	7.2	-5.7	-49.6

Impact on estimated antenna heights: changed datum station set



Except for WETTZ13N, the median changes in height have the expected sign.

> The most significant impact is given for WETTZELL (which has the largest maximum excess delay) and KOKEE.

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Impact on estimated antenna heights: unchanged datum station set



- > Basically all session-wise changes in height have the expected sign.
- > The median changes are similar while the variation of changes is much smaller compared to the previous case.

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Datum (scale) parameter time series: ITRF2014 and GD

solution	a priori TRF	gravitational deformation (GD)	datum stations
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dgf2023a old GD	ITRF2020	first 6 antennas	all
dgf2023a	ITRF2020	first 6 + 7 new antennas	all except for 7 new GD antennas

- The figure shows the moving 20 session median of the scale parameter of 7 parameter Helmert transformations of the session-wise solutions w.r.t. their a priori TRFs (outliers: |scale| > 60 mm).
- Until about 2015, the median scales w.r.t. ITRF2014 and ITRF2020 are very similar.
- There is no continuous scale drift between dgf2020a and the ITRF2014 (blue) afterwards.
- The 7 new gravitational deformation models seem to have no significant impact on the scale (yellow).



Conclusions

- The most significant change with the secular ITRS 2020 realizations is given for stations with a small observation history before 2015.
- The impact of switching the a priori TRF in VLBI analysis is more pronounced if the set of datum stations is also altered.
- A drift is observed in the scale w.r.t. the ITRF2020 and apparently also w.r.t. the JTRF2020, starting shortly before 2015 as well.
- The scale drift is not observed w.r.t. DTRF2020 (are there systematic differences in up direction, or is the choice of techniques for realizing the scale, i.e., SLR vs. GNSS, relevant?) and ITRF2014.
- The impact of the new gravitational deformation models on the estimated heights of the corresponding antennas is as expected on average, but it also depends on the chosen set of datum stations.
- The scale is hardly affected by the new gravitational deformation models, which have comparatively small excess delays.